

Superconductivity in solid benzene molecular crystal

HPSTAR
564-2018

Guo-Hua Zhong¹ , Chun-Lei Yang¹, Xiao-Jia Chen² and Hai-Qing Lin³

¹ Shenzhen Institutes of Advanced Technology, Chinese Academy of Sciences, Shenzhen 518055, People's Republic of China

² Center for High Pressure Science and Technology Advanced Research, Shanghai 201203, People's Republic of China

³ Beijing Computational Science Research Center, Beijing 100193, People's Republic of China

E-mail: xjchen@hpstar.ac.cn and haiqing0@csrc.ac.cn

Received 19 March 2018, revised 1 May 2018

Accepted for publication 11 May 2018

Published 23 May 2018



Abstract

Light-element compounds hold great promise of high critical temperature superconductivity judging from the theoretical perspective. A hydrogen-rich material, benzene, is such a kind of candidate but also an organic compound. A series of first-principles calculations are performed on the electronic structures, dynamics properties, and electron–phonon interactions of solid benzene at high pressures. Benzene is found to be dynamically stable in the pressure range of 180–200 GPa and to exhibit superconductivity with a maximum transition temperature of 20 K at 195 GPa. The phonon modes of carbon atoms are identified to mainly contribute to the electron–phonon interactions driving this superconductivity. The predicted superconductivity in this simplest pristine hydrocarbon shows a common feature in aromatic hydrocarbons and also makes it a bridge to organic and hydrogen-rich superconductors.

Keywords: benzene, superconductivity, pressure, hydride, hydrocarbon

(Some figures may appear in colour only in the online journal)

1. Introduction

Exploring new materials with high critical temperature (T_c) superconductivity is one attractive field in modern condensed matter physics. Although there has been no specific method to guide the search for superconductors, the light elements or their bearing compounds such as hydrogen and hydrogen-rich materials are believed to be promising candidates. The basis of the latter is mainly from Ashcroft's theoretical viewpoint that hydrogen dominant metallic alloys can reduce the metallization pressure to the scope of the experiment and become the potential candidates of high- T_c superconductors [1]. In 2008, Chen *et al* [2] observed the signature of metallization of solid silane (SiH_4) at pressure above 60 GPa. Subsequently, Eremets *et al* [3] measured superconductivity with $T_c \sim 17$ K in solid SiH_4 , though there are still some queries about these findings. These experimental observations imply the feasibility of Ashcroft's theory. Recently, superconductivity at around $T_c \sim 200$ K was experimentally observed in the sulfur hydride system [4, 5], illustrating the feasibility of seeking for

the high- T_c superconductors in hydrogen-rich materials. Prior to this discovery, Duan *et al* [6] predicted the H_3S ($Im\bar{3}m$) structure for sulfur hydride at high pressure. These authors predicted potential superconductivity with high T_c value of 191–204 K at 200 GPa [6]. The predicted crystal structure was soon confirmed in experiments [7–9], though many other sulfur hydrides form such as H_4S_3 , H_5S_8 , H_3S_5 , and HS_2 . It has also been suggested [6] that the electron–phonon coupling mainly arises from H vibrations in this hydride. For instance, the contribution of H vibrations to the coupling reaches to 90% in the sulfur hydride system. These results demonstrate the extreme importance that pressure played in the discovery of superconductivity of light materials as well as in hybridizing interaction of H with other elements.

In addition, organic based compounds were also suggested as candidates of high temperature or room temperature superconductors [10, 11]. This idea assumed that the interaction of electrons with much higher excitation energy than the phonon energy can result in a substantially higher T_c . The first experimental evidence of superconductivity in organic metals was

found in 1980 [12]. Since then, numerous organic superconductors have been reported including electron donor and electron acceptor molecules. Superconductivity with T_c as high as 38 K was observed in cesium-doped fullerene [13], and more than 120 K was also found in potassium-doped *p*-terphenyl [14]. These findings highlight that organic compounds have potential to become high- T_c superconductors. This is mainly because non phonon mechanisms were found to account for their superconductivity in these kinds of materials. Only electron–phonon coupling is not enough to produce such high- T_c superconductivity in alkali metal doped fullerenes [13] and aromatic hydrocarbons [15]. Meanwhile, the low dimensional feature of organic molecule was proposed to favor the strong electronic correlation effects in these materials such as doped fullerenes [16] and aromatic hydrocarbons [17–19]. While pressure can greatly enhance superconductivity in aromatic hydrocarbons such as potassium-doped phenanthrene [20] and picene [21], the realization of superconductivity solely by applying pressure on pristine organic compounds has not been achieved yet.

In this work, we choose benzene (C_6H_6), a hybrid of the simplest aromatic hydrocarbon and a hydrogen-rich material, to explore superconductivity at high pressures. Solid C_6H_6 was previously predicted to enter a metallic state at pressures above 180 GPa [22], though the metallization had not been realized experimentally [23]. Through extensive calculations of electronic structures, dynamical properties and electron–phonon interactions, we find that C_6H_6 is in fact superconducting in the pressure range of 180 and 200 GPa with a maximum T_c of 20 K. Within the framework of electron–phonon coupling, the superconductivity is examined to mainly come from the contribution of C element, differing from the H dominant materials. Based on our systematical investigations on the superconductivity of potassium-doped aromatic hydrocarbons in recent years [24], we conclude that the materials containing benzene rings must be superconducting, with $T_c \sim 5\text{--}7$ K, and the superconductivity is increased with the change of structure, electronic correlations and pressure.

2. Computational methods

To study the structural and electronic properties of solid C_6H_6 , we employed the Vienna *ab initio* simulation package (VASP) [25, 26] based on the projector augmented wave method. For the plane-wave basis-set expansion, an energy cutoff of 800 eV was adopted. Dense *k*-point meshes were used to sample the first Brillouin zone and ensured that energies converged to within 1 meV/atom. At the same time, lattice dynamics and electron–phonon interactions were calculated using density functional perturbation theory [27] and the Troullier–Martins norm-conserving potentials [28], as implemented in the QUANTUM-ESPRESSO (QE) code [29]. The cutoff energies of 80 and 600 Ry were used for wave functions and charge densities, respectively. $24 \times 24 \times 24$ Monkhorst–Pack *k*-point grid with Gaussian smearing of 0.003 Ry was used for the electron–phonon interaction matrix element calculation at $6 \times 6 \times 6$ *q*-point mesh. In both VASP and QE codes, the

local density approximation (LDA) [30, 31] functional was selected. Forces and stresses for the converged structures were optimized and checked to be within the error allowance of the VASP and QE codes. The computational methods have been proved to be reliable in previous reports [6, 22].

3. Results and discussion

For solid C_6H_6 , it has well-known that the pressure can change its morphology and structure and lead to a series of phase transitions [32–34]. At 1.4 GPa, the transition occurs from *Pbca* phase to $P4_32_12$ phase. At 4 GPa, phase II transfers to $P2_1/c$ phase, and up to 11 GPa. However, at higher pressure above 11 GPa, the experimental results are still not perfect. Keeping the molecular characteristic of benzene, the previous theoretical study pointed out that [22], solid benzene transforms to $P2_1$ from $P2_1/c$ at 40 GPa and remains to 300 GPa. Noticeably, from the energy point of view, the graphane-like crystal possesses the lower enthalpy than molecular benzene crystal starting from about 10 GPa to 300 GPa. $P2_1$ and $P2_1/c$ are only two metastable phases in the high pressure region [22]. However, it is not easy to convert benzene to polymer or amorphization compound. Because there are likely significant barriers among these interconverting processes and the conversion reaction also involves temperature condition except for pressure [22, 35]. Thus, both $P2_1$ and $P2_1/c$ phases can keep the feature of benzene molecule instead of the amorphous structure of C and H at high pressure. Especially, $P2_1/c$ phase of C_6H_6 can behave as a metal in the pressure range of 180–200 GPa [22].

As a comparison and a check on accuracy, we start our investigation by looking at the first solid phase of benzene with *Pbca* symmetry. Our optimized crystal constants are respectively $a = 7.041 \text{ \AA}$, $b = 8.903 \text{ \AA}$ and $c = 6.357 \text{ \AA}$ at zero temperature, which are 3%–6% less than experimental values at 78 K [36]. If the effect of temperature is considered, the error between theoretical prediction and experimental measurement is acceptable. Moreover, for *Pbca* C_6H_6 , our calculated band gap of 4.1 eV is in a good agreement with previous results [22]. Based on the test, we have optimized the crystal structures of $P2_1/c$ C_6H_6 in the range of 180–200 GPa. Figure 1 shows the geometrical structure of $P2_1/c$ phase of C_6H_6 crystal. In the case of 190 GPa, the optimized lattice constants $a = 3.962 \text{ \AA}$, $b = 3.881 \text{ \AA}$ and $c = 5.287 \text{ \AA}$ as well as the angle $\beta = 100.3^\circ$.

The calculated band gap of 4.1 eV for C_6H_6 is far less than 7.5 eV of methane (CH_4) [37], which indicates C_6H_6 is more easily to become into metal under pressure comparing with CH_4 . Indeed, the band gap of C_6H_6 with $P2_1/c$ structure has been closed when the pressure increases to 180 GPa, and keeping the metallic behavior up to 200 GPa. On the contrary, until 520 GPa, CH_4 is still a semiconductor [38]. In the case of 190 GPa, we show the electronic band structures, density of states (DOS) and Fermi surface sheets of C_6H_6 in figure 2. As shown in figure 2(a), similar to other hydrogen-rich materials, the metallization of C_6H_6 are mainly derived from the increase of covalent interaction under pressure. However, the nature

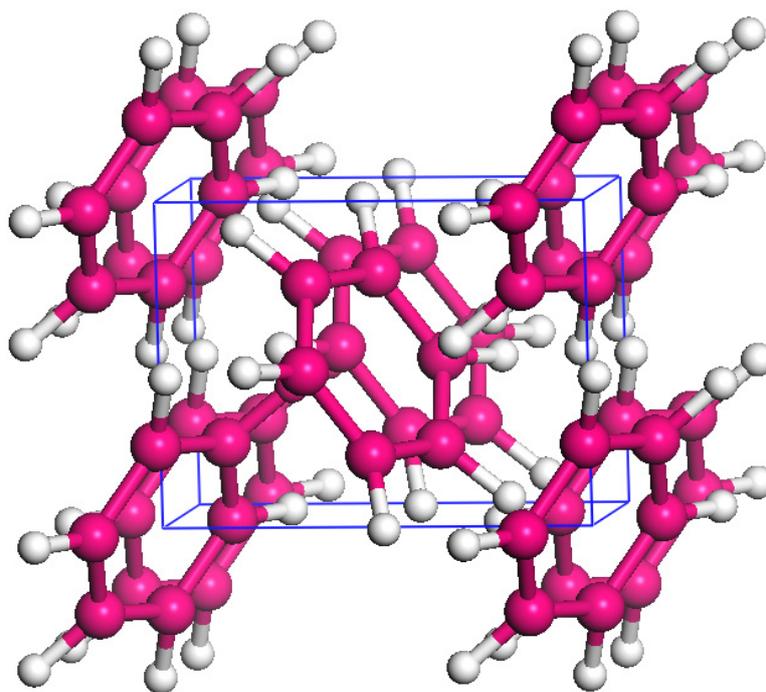


Figure 1. The geometrical structure of C_6H_6 crystal in the high pressure range of 180–200 GPa. Red and white spheres represent C and H atoms, respectively.

of energy bands of C_6H_6 is different from those of doped aromatic hydrocarbons where the band structure possessed the typical charge transfer characteristics. Corresponding to band structure, a small amount of electronic states gathers at Fermi level in the DOS picture, which is mainly contributed by C-2 p states (figure 2(b)). The DOS at Fermi level is about 0.32 states/eV, which is a small value comparing with those of doped aromatic hydrocarbons. But, we find that the DOS at Fermi level continuously increases with the increase of pressure in this range of 180–200 GPa. Checking the fine electronic feature near Fermi level, we find additionally that the VB1 and VB2 bands (marked in figure 2(a)) form the hole-like Fermi surfaces around B k -point, as shown in figure 2(c). The CB1 band with higher energy forms the electron-like Fermi surface around Γ k -point.

For this metallic phase of C_6H_6 , our calculated phonon dispersion (figure 3(a)) has confirmed the dynamical stability of $P2_1/c$ structure due to the absence of imaginary frequency modes. The visible gaps in the phonon spectra divide phonon frequency into three parts, low-frequency below 22.7 THz, intermediate-frequency from 24.3 to 56.0 THz and high-frequency above 102.0 THz. Combining with the projected phonon density of states shown in figure 3(b), we can determine that low-frequency modes come mainly from the C atomic motions, the intermediate-frequency bands are induced by the organic intermolecular coupling, while the high-frequency spectrum belongs to the C–H bond stretching vibrations. Corresponding to the origin of phonon modes, the calculated results show that the higher pressure leads to the slightly hardening of intermediate/high-frequency phonon and the softening of low-frequency phonon.

Further understanding the dynamical properties of C_6H_6 under pressure, we compared it with typical hydrogen-rich

materials and organic compounds. For SiH_4 , the material exists in form of multifold-coordinated silicon hydride at 200 GPa [39]. The low-frequency induced by the motions of Si is below 19 THz, the intermediate-frequency region is between 19 and 33 THz, dominated by intermolecular (Si–H–Si) interactions, while the high-frequency phonon from Si–H stretching vibrations is in the range of 33–75 THz. This implies stronger stretching vibration of C–H bond in C_6H_6 than Si–H bond in SiH_4 . Between C_6H_6 and SiH_4 , their regions of low-frequency from the heavier element are almost accordant, but there is big difference for the intermediate- and high-frequency modes. For germane (GeH_4), the multifold-coordinated germanium hydride and the quasi H_2 molecules coexist in the system at 220 GPa [40]. The phonon frequency of GeH_4 is also obviously divided into three parts similar to C_6H_6 , due to the existence of gaps in phonon energy. The vibration modes of heavier element Ge is below 14 THz, the intermediate-frequency region is between 14 and 62 THz owing to the intermolecular H_2 coupling and the Ge–H stretching vibrations, while the high-frequency phonon from the intramolecular H_2 vibrations is in the region of 73–80 THz. Comparing with GeH_4 , both phonon vibrations and intermolecular coupling in C_6H_6 are stronger. In H_2 -containing compounds such as $PbH_4(H_2)_2$, the alloy like compound of Pb and quasi H_2 units is formed at 200 GPa [41]. The low-frequency phonon below 8.6 THz comes from heavier element Pb. But the region of intermediate-frequency vibrations in 8.6–56 THz arises from the intermolecular H_2 coupling. Two regions of high-frequency in ranges of 81–87 THz and 96–101 THz are induced by the intramolecular vibrations of two kinds of quasi H_2 molecules. Comparing with $PbH_4(H_2)_2$, C_6H_6 has almost the same intermediate-frequency phonon coupling, but stronger low-frequency vibrations. For sulfur hydride, H_3S is

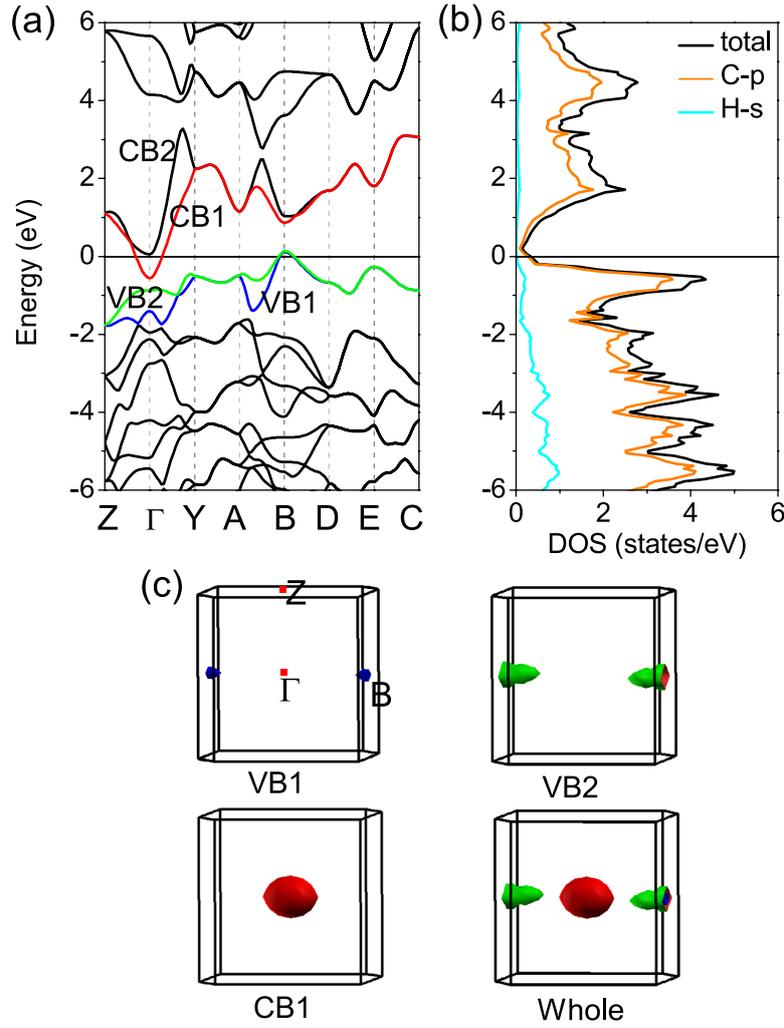


Figure 2. Electronic structures of C_6H_6 at 190 GPa: (a) energy band structure, (b) total and projected density of states (DOS) and (c) Fermi surfaces. Zero energy denotes the Fermi level.

the main form at 200 GPa with $Im\bar{3}m$ space-group [6]. At this pressure, the vibration modes below 18 THz are due to the motions of S atom, the other phonon bands between 18 and 55 THz are formed by the S–H bond stretching vibrations. There is no higher phonon frequency than 55 THz in H_3S system. In addition, for superconducting aromatic hydrocarbons, in the K_3 picene [42], below 9 THz, the vibration modes come from the motions of potassium as well as the coupling between it and organic molecules, the intermolecular coupling vibrations appear in the range of 9–47 THz, and the high-frequency C–H stretching modes is localized around 91 THz. From this comparison, there is stronger intermolecular coupling vibrations in C_6H_6 at pressures than K_3 picene.

Based on electronic and phonon properties above, the electron–phonon coupling $\lambda(\omega)$, logarithmic average phonon frequency ω_{\log} and the Eliashberg phonon spectral function $\alpha^2F(\omega)$ have been investigated to explore the possible superconductivity of C_6H_6 . The Eliashberg function was calculated for each phonon mode ν with wavevector \mathbf{q} by the following equation

$$\alpha^2F(\omega) = \frac{1}{2\pi N(\varepsilon_F)} \sum_{\mathbf{q}\nu} \frac{\gamma_{\mathbf{q}\nu}}{\omega_{\mathbf{q}\nu}} \delta(\omega - \omega_{\mathbf{q}\nu}), \quad (1)$$

where

$$\gamma_{\mathbf{q}\nu} = 2\pi\omega_{\mathbf{q}\nu} \sum_{mn} \sum_{\mathbf{k}} |g_{\mathbf{k}+\mathbf{q},\mathbf{k}}^{\mathbf{q}\nu, mn}|^2 \delta(\varepsilon_{\mathbf{k}+\mathbf{q},m} - \varepsilon_F) \delta(\varepsilon_{\mathbf{k},n} - \varepsilon_F). \quad (2)$$

Correspondingly, the electron–phonon coupling was also calculated by the following equation

$$\lambda_{\mathbf{q}\nu} = \frac{\gamma_{\mathbf{q}\nu}}{\pi N(\varepsilon_F) \omega_{\mathbf{q}\nu}^2}. \quad (3)$$

Referring to the report by Casula *et al* [42], we decomposed the $\alpha^2F(\omega)$ by projecting on C and H phonons, plotted in figure 3(c), and the correspondingly integral $\lambda(\omega) = 2 \int_0^\omega d\omega' \alpha^2F(\omega')/\omega'$, shown in figure 3(d), which gives the total $\lambda = \sum_{\mathbf{q}\nu} \lambda_{\mathbf{q}\nu}$ in the $\omega \rightarrow \infty$ limit. From the $\alpha^2F(\omega)$ and $\lambda(\omega)$ shown in figure 3, the total electron–phonon coupling λ is mainly contributed by the low-frequency phonon from C atom, namely the motion of electrons on benzene ring plane. The rest small account of total λ comes from the organic intermolecular coupling. In the case of 190 GPa of $P2_1/c$ structure, the total λ is 0.68, a moderate-intensity electron–phonon coupling. The low-frequency vibration modes from C atoms contribute 83% of total λ , while the organic intermolecular coupling has only 17% contribution to total

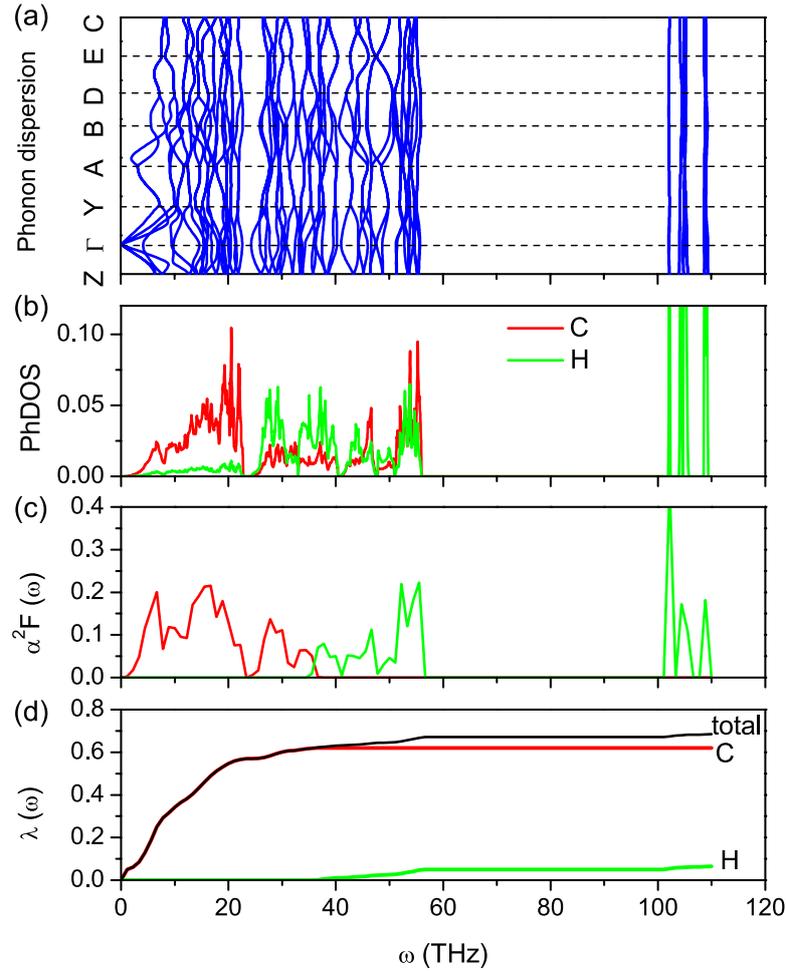


Figure 3. Calculated phonon dispersion curves (a), phonon density of states (b), Eliashberg spectral function $\alpha^2F(\omega)$ (c) and electron–phonon coupling integral $\lambda(\omega)$ resolved by projections on C and H phonon (d) for $P2_1/c$ structure of benzene at 190 GPa.

λ , and the high-frequency phonon from C–H bond stretching hardly couples the electron. Additionally, we have examined the pressure effect for phonon frequency, phonon density of states, Eliashberg function and electron–phonon coupling, shown in figure 4. We can find the obvious softening of phonon around A point (figure 4(a)) and the wider phonon band (figure 4(b)) and Eliashberg spectrum (figure 4(c)) with the increase of pressure. Thus, the total λ changes big by increasing pressure, from 0.53 at 185 GPa to 0.85 at 195 GPa as shown in figure 4(d).

Adopting the modified McMillan equation by Allen and Dynes [43] and combining with the obtained λ and ω_{\log} , we analyzed the dependence of the superconducting critical temperature of C_6H_6 on pressure, shown in figure 5. However, ω_{\log} decreases with increasing pressure. On the other hand, λ is enhanced by the applied pressure. As a result, we obtain a slight increase for T_c in the pressure range of 180–200 GPa by using commonly accepted values of the Coulomb pseudo-potential $\mu^* = 0.1$ –0.13. As shown in figure 5, at 195 GPa, T_c reaches 19.2 K in C_6H_6 for $\mu^* = 0.1$. Thus, we theoretically suggest that benzene is superconducting at high pressure, with the T_c close to 20 K.

By analyzing superconductivity, we find that the total λ of $K_3picene$ [42] is mainly due to the low-frequency phonon

coupling to electron from potassium, and their λ values, 0.53–0.85 in the range of 185–195 GPa for C_6H_6 and 0.73 for $K_3picene$, are comparable in quantity. The difference is that C_6H_6 has higher ω_{\log} of 360.3–720.4 K than $K_3C_{22}H_{14}$. Hence, the predicted T_c of C_6H_6 is higher than that of $K_3picene$ [42]. Comparing with hydrogen-rich materials, the same is that the pressure induced the metallization, but the difference is that the total λ of hydrogen dominate material is mainly contributed by the intermediate-frequency phonon coupling electron. The higher λ is often obtained in hydrogen-rich systems. Especially, the H_3S results in the λ of 2.19 at 200 GPa. Except for the small λ value, the ω_{\log} of C_6H_6 is less than half of H_3S . As a result, the T_c of C_6H_6 is much less than that of H_3S . Combining with the comparison of phonon frequency above, we know that the low and intermediate-frequency phonon coupling electron mainly contributes to the superconductivity. On one hand, the phonon softening in this region will be help for the improvement of the T_c in C_6H_6 . On the other hand, the DOS at Fermi level is only 0.12–0.17 states/ev/f.u. for C_6H_6 in the range of 185–195 GPa. This value is much less than 0.75 states/ev/f.u. for $PbH_4(H_2)_2$ at 200 GPa [41] and ~ 0.4 states/ev/f.u. at 200 GPa [6] (or 0.51 states/ev/f.u. at 210 GPa [44]) for H_3S . Therefore, more Cooper pairing electrons are

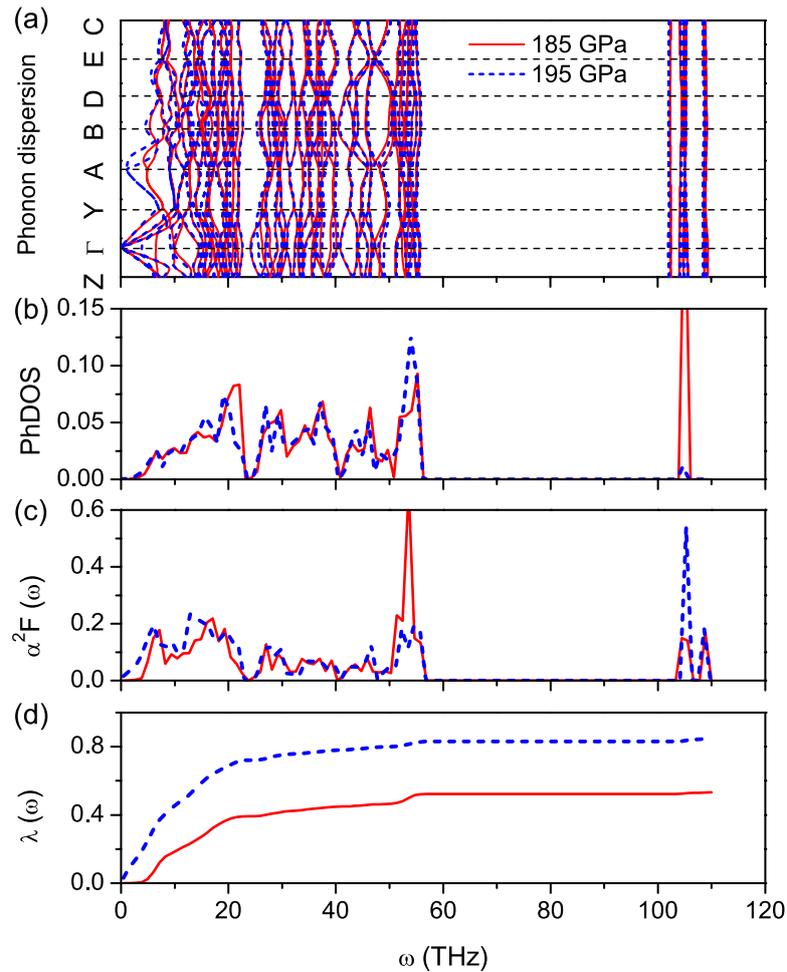


Figure 4. Calculated phonon dispersion curves (a), phonon density of states (b), Eliashberg spectral function $\alpha^2F(\omega)$ (c) and electron-phonon coupling integral $\lambda(\omega)$ (d) for $P2_1/c$ structure of benzene at 185 GPa (solid line) and 195 GPa (dashed line).

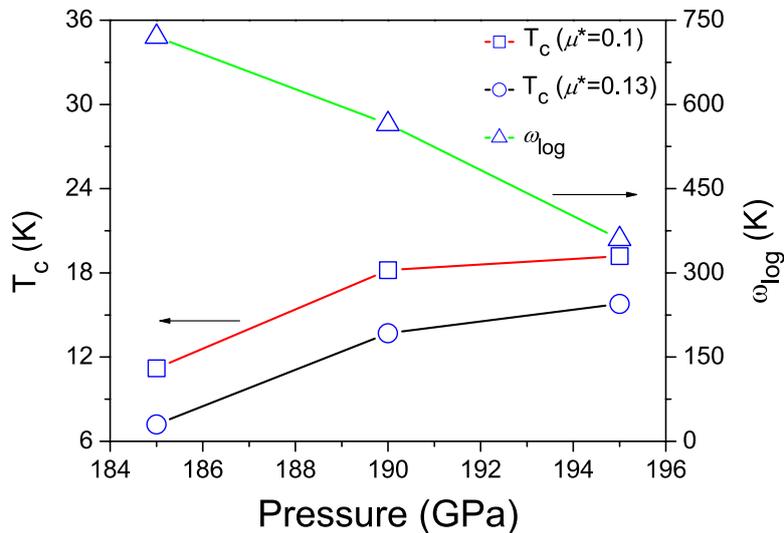


Figure 5. Pressure dependence of the superconducting critical temperature T_c and the logarithmic average phonon frequency ω_{log} for benzene in $P2_1/c$ phase.

another key to achieve the stronger electron-phonon coupling in C_6H_6 . For instance, pressure leads to the increase of DOS at Fermi level, T_c rises with increasing pressure in C_6H_6 , accordingly.

4. Conclusions

In conclusion, with the aim of exploring superconductivity in a compound containing only carbon and hydrogen, we

have studied electronic structures, dynamics properties, and electron–phonon interactions of C₆H₆ at high pressures. We revealed that C₆H₆ is superconducting in the pressure range of 180–200 GPa, which is further evidence that the system containing benzene rings must be superconducting. T_c was found to gradually increase with the increase of pressure, reaching 20 K at 195 GPa. Differing from the H-dominated materials, the phonon vibrations of C atom mainly contribute to the electron–phonon coupling. The T_c higher than that of K-doped benzene, phenanthrene and picene mainly comes from the larger ω_{\log} induced by pressure. The prediction of superconductivity in C₆H₆ under pressure will call for experimental testing and the comparison of effects from the electron–phonon coupling and the electronic correlation in such a light element material.

Acknowledgments

The work was supported by the National Natural Science Foundation of China (Grant Nos. 11274335, 61574157 and 61774164) and the Basic Research Program of Shenzhen (JCYJ20170818153404696, JCYJ20160331193059332, JCYJ20150925163313898 and JCYJ20160331193134437). The partial calculation was supported by the Special Program for Applied Research on Super Computation of the NSFC-Guangdong Joint Fund (the second phase) under Grant No. U1501501.

ORCID iDs

Guo-Hua Zhong  orcid.org/0000-0003-0673-8738

References

- [1] Ashcroft N W 2004 *Phys. Rev. Lett.* **92** 187002
- [2] Chen X J, Struzhkin V V, Song Y, Goncharov A F, Ahart M, Liu Z, Mao H K and Hemley R J 2008 *Proc. Natl Acad. Sci. USA* **105** 20–3
- [3] Eremets M I, Trojan I A, Medvedev S A, Tse J S and Yao Y 2008 *Science* **319** 1506–9
- [4] Drozdov A P, Eremets M I, Troyan I A, Ksenofontov V and Shylin S I 2015 *Nature* **525** 73–6
- [5] Troyan I, Gavriluk A, Ruffer R, Chumakov A, Mironovich A, Lyubutin I, Perekalin D, Drozdov A P and Eremets M I 2016 *Science* **351** 1303–6
- [6] Duan D, Liu Y, Tian F, Li D, Huang X, Zhao Z, Yu H, Liu B, Tian W and Cui T 2014 *Sci. Rep.* **4** 6968
- [7] Einaga M, Sakata M, Ishikawa T, Shimizu K, Eremets M I, Drozdov A P, Troyan I A, Hirao N and Ohishi Y 2016 *Nat. Phys.* **12** 835–8
- [8] Li Y *et al* 2016 *Phys. Rev. B* **93** 020103
- [9] Goncharov A F, Lobanov S S, Kruglov I, Zhao X M, Chen X J, Oganov A R, Konôpková Z and Prakapenka V B 2016 *Phys. Rev. B* **93** 174105
- [10] Little W A 1964 *Phys. Rev.* **134** A1416–24
- [11] Ginzburg V L 1976 *Sov. Phys.—Usp.* **19** 174–9
- [12] Jerome D, Mazaud A, Ribault M and Bechgaard K 1980 *J. Phys. Lett.* **41** L95–8
- [13] Ganin A Y, Takabayashi Y, Khimyak Y Z, Margadonna S, Tamai A, Rosseinsky M J and Prassides K 2008 *Nat. Mater.* **7** 367–71
- [14] Wang R S, Gao Y, Huang Z B and Chen X J 2017 Superconductivity above 120 kelvin in a chain link molecule (arXiv:1703.06641)
- [15] Mitsuhashi R *et al* 2010 *Nature* **464** 76–9
- [16] Ganin A Y *et al* 2010 *Nature* **466** 221–5
- [17] Huang Z, Zhang C and Lin H Q 2012 *Sci. Rep.* **2** 922
- [18] Giovannetti G and Capone M 2011 *Phys. Rev. B* **83** 134508
- [19] Kim M, Min B I, Lee G, Kwon H J, Rhee Y M and Shim J H 2011 *Phys. Rev. B* **83** 214510
- [20] Wang X F, Liu R H, Gui Z, Xie Y L, Yan Y J, Ying J J, Luo X G and Chen X H 2011 *Nat. Commun.* **2** 507
- [21] Kubozono Y *et al* 2015 *Physica C* **514** 199–205
- [22] Wen X D, Hoffmann R and Ashcroft N W 2011 *J. Am. Chem. Soc.* **133** 9023–35
- [23] Hillier N J and Schilling J S 2014 *High Press. Res.* **34** 1–8
- [24] Zhong G H, Chen X J and Lin H Q 2015 Prediction of superconductivity in potassium-doped benzene (arXiv:1501.00240)
- [25] Kresse G and Furthmuller J 1996 *Comput. Mater. Sci.* **6** 15–50
- [26] Kresse G and Furthmuller J 1996 *Phys. Rev. B* **54** 11169–86
- [27] Baroni S, de Gironcoli S, Dal Corso A and Giannozzi P 2001 *Rev. Mod. Phys.* **73** 515–62
- [28] Troullier N and Martins J L 1991 *Phys. Rev. B* **43** 1993–2006
- [29] Giannozzi P *et al* 2009 *J. Phys.: Condens. Matter* **21** 395502
- [30] Ceperley D M and Alder B J 1980 *Phys. Rev. Lett.* **45** 566–9
- [31] Perdew J P and Zunger A 1981 *Phys. Rev. B* **23** 5048–79
- [32] Thiéry M M and Léger J M 1988 *J. Chem. Phys.* **89** 4255–71
- [33] Ciabini L, Gorelli F A, Santoro M, Bini R, Schettino V and Mezouar M 2005 *Phys. Rev. B* **72** 094108
- [34] Ciabini L, Santoro M, Gorelli F A, Bini R, Schettino V and Raugei S 2007 *Nat. Mater.* **6** 39–43
- [35] Hou K C and Palmer H B 1965 *J. Phys. Chem.* **69** 863–8
- [36] Kozhin V M and Kitaigorodskii A I 1955 *Zh. Fiz. Khim.* **29** 2074–5
- [37] Lin H, Li Y L, Zeng Z, Chen X J and Lin H Q 2011 *J. Chem. Phys.* **134** 064515
- [38] Martinez-Canales M and Bergara A 2006 *High Press. Res.* **26** 369–75
- [39] Chen X J, Wang J L, Struzhkin V V, Mao H K, Hemley R J and Lin H Q 2008 *Phys. Rev. Lett.* **101** 077002
- [40] Gao G, Oganov A R, Bergara A, Martinez-Canales M, Cui T, Iitaka T, Ma Y and Zou G 2008 *Phys. Rev. Lett.* **101** 107002
- [41] Cheng Y, Zhang C, Wang T, Zhong G, Yang C, Chen X J and Lin H Q 2015 *Sci. Rep.* **5** 16475
- [42] Casula M, Calandra M, Profeta G and Mauri F 2011 *Phys. Rev. Lett.* **107** 137006
- [43] Allen P B and Dynes R C 1975 *Phys. Rev. B* **12** 905–22
- [44] Papaconstantopoulos D A, Klein B M, Mehl M J and Pickett W E 2015 *Phys. Rev. B* **91** 184511